

Relativistic MHD

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§ 0. Prelude. Metric tensor, inertial frame, etc.

0.1. General stuff

$$ds^2 = g_{\nu\mu} dx^\nu dx^\mu \quad (0.1)$$

- spacetime metric form.

$$(\nu, \mu = 0, 1, 2, 3)$$

$\{x^\nu\}$ - spacetime coordinates

• The coordinate basis of vectors:

$$\vec{e}_\nu = \frac{\partial}{\partial x^\nu} \quad (0.2)$$

$$g_{\nu\mu} = (\vec{e}_\nu \cdot \vec{e}_\mu) \quad (0.3)$$

• Normalized basis:

$$\hat{\vec{e}}_\nu = \frac{1}{\sqrt{|g_{\nu\nu}|}} \vec{e}_\nu \quad (0.4)$$

$$\Rightarrow |\hat{\vec{e}}_\nu|^2 = \pm 1 \quad (0.5)$$

• Components of vectors

Contravariant, $\{A^\nu\}$: $\vec{A} = A^\nu \vec{e}_\nu \quad (0.6)$

(Notice the use of Einstein's summation rule.)

As a rule we will ~~not~~ use normalized basis in what follows

Covariant, $\{A_\nu\}$: $A_\nu = g_{\nu\mu} A^\mu$ (0.7)

0.2. 3+1 splitting of inertial frames

(In curved spacetime only locally inertial frames)

The ~~special~~ ^{static} coordinate grid is fixed in space. (~~static~~ grid)! No deformations, rotation, etc.

$ds^2 = -c^2 dt^2 + dl^2$ (0.8)

$dl^2 = \delta_{ij} dx^i dx^j$ (0.9)

δ_{ij} - metric tensor of space

$\{x^i\}$, $i=1,2,3$ - the spatial coordinates

$\delta_{ij} = (\underline{e}_i \cdot \underline{e}_j)$ (0.10)

$\underline{e}_i = \frac{\partial}{\partial x^i}$ - basis vectors of spatial coordinates

$x^0 = ct$ (0.11)

- time coordinate of spacetime

Thus, $ds^2 = -dx^0^2 + \delta_{ij} dx^i dx^j$

and

$g_{\nu\mu} = \begin{pmatrix} -1 & 0 & 0 & 0 \\ 0 & & & \\ 0 & & \delta_{ij} & \\ 0 & & & \end{pmatrix}$ (0.12)

- metric tensor of inertial frame

$$g^{\nu\mu} = \begin{pmatrix} -1 & 0 & 0 & 0 \\ 0 & & & \\ 0 & & & \\ 0 & & & (\gamma^{ij}) \end{pmatrix} \quad (0.13)$$

where (γ^{ij}) is inverse to (γ_{ij}) ,
 that is $\gamma^{ik} \gamma_{kj} = \delta^i_j$.

$$g \equiv \det(g_{\nu\mu}) = -\gamma \quad (0.14)$$

$$\gamma \equiv \det(\gamma_{ij}) \quad (0.15)$$

Note that

$$\frac{\partial \gamma_{ij}}{\partial t} = 0, \quad \frac{\partial g}{\partial t} = 0 \quad (0.16)$$

- ~~static~~ special grid ?

• Components of vectors and tensors

$$A^\nu = (\alpha, a^1, a^2, a^3) \equiv (\alpha, a^i) \quad (0.17)$$

where α is a special scalar (3-scalar)

\underline{a} is a special vector (3-vector)

$$A_\nu = (-\alpha, a_i), \quad (0.18)$$

where $a_i = \gamma_{ij} a^j$.

For any second rank tensor

$$\Pi^{\nu\mu} = \begin{pmatrix} a & b^j \\ c^i & \pi^{ij} \end{pmatrix} \quad (0.19)$$

a - 3-scalar, \vec{b}, \vec{c} - 3-vectors, π^{ij} - 3-tensor

$$\Pi_{\nu\mu} = \begin{pmatrix} a & -b_i \\ -c_j & \Pi_{ij} \end{pmatrix} \quad (0.20)$$

where $\Pi_{ij} = \delta_{ia} \delta_{jb} \Pi^{ab}$ are the covariant components of Π .

E.T.C.

• Example of proof calculation

$$\begin{aligned} \Pi_{0j} &= g_{0\nu} g_{j\mu} \Pi^{\nu\mu} = g_{00} g_{j\mu} \Pi^{0\mu} = \\ &= g_{00} g_{jk} \Pi^{0k} = -\delta_{jk} b^k = -b_j \end{aligned}$$



§1 Particle Dynamics

$$p^\nu = m_0 u^\nu \quad (1.1)$$

- energy-momentum vector (4-momentum).

m_0 - rest mass of a particle

$$u^\nu = \frac{dx^\nu}{d\tau} \quad (1.2) \quad \text{- 4-velocity}$$

τ - proper time of the particle

• Energy-momentum conservation law:

$$\sum_{k=1}^N p_{(k)}^\nu = \text{const} \quad (1.3)$$

(N particles)

• Inertial frame 3+1 splitting

$$u^\nu = (\Gamma c, \Gamma v^i) \quad (1.4)$$

$$\Gamma = \frac{dt}{d\tau} \quad \text{- Lorentz factor} \quad (1.5)$$

$$v^i = \frac{dx^i}{dt} \quad (1.6) \quad \text{-}$$

$$p^\nu = (m_0 \Gamma c, m_0 \Gamma v^i) = \left(\frac{E}{c}, \underline{P} \right) \quad (1.7)$$

where

$$E = m c^2 \quad \text{- energy} \quad (1.8)$$

$$m = \Gamma m_0 \quad \text{- inertial mass} \quad (1.9)$$

$$\underline{P} = m \underline{v} \quad \text{- 3-momentum} \quad (1.10)$$

The minimum energy

$$E_0 = m_0 c^2 \quad \text{— rest mass-energy} \quad (1.11)$$

$$K = E - E_0 = m_0 (\gamma - 1) c^2 \quad (1.12)$$

— kinetic energy

$$m = m_0 + m_0 (\gamma - 1) = m_0 + \frac{K}{c^2} \quad (1.13)$$

- (i) The rest mass-energy contributes to the total energy!
- (ii) The kinetic energy contributes to particles' inertial mass!

For a gas element these mean that

- (i) Its energy includes the rest mass of its particles.
- (ii) Its inertial mass includes the gas thermal energy.

SS2 Thermodynamics

2.1

It describes the properties of gas in the gas frame.

All these quantities are measured in the gas frame:

n - volume number density of particles

$\rho = m_0 n$ - volume rest-mass density

$P = nkT$ - gas pressure

T - temperature.

$$\boxed{\mathcal{E} = \rho c^2 + \mathcal{E}_t} \quad (2.1)$$

- volume energy density.

\mathcal{E}_t - thermal energy density

$$\boxed{\omega = \mathcal{E} + P} \quad (2.2)$$

- enthalpy

• Polytropic equation of state (EOS)

$$\boxed{\mathcal{E}_t = \frac{1}{\tilde{\gamma} - 1} P} \quad (2.3)$$

$$\boxed{\omega = \rho c^2 + \frac{\tilde{\gamma}}{\tilde{\gamma} - 1} P} \quad (2.4)$$

$1 < \tilde{\gamma} < 2$ - ratio of specific heats

§53 Hydrodynamics

3.1 The continuity equation

(The rest mass conservation)

Not always satisfied.

• Nonrelativistic case

$$\boxed{\frac{\partial \rho}{\partial t} + \nabla_i \rho v^i = 0} \quad (3.1)$$

∇_i - covariant derivative

$$\nabla_i a^j = \frac{\partial a^j}{\partial x^i} - \Gamma^j_{ik} a^k \quad (3.2)$$

where Γ^j_{ik} - Cristoffel's symbols

For divergence this can be simplified!

$$\boxed{\nabla_i a^i = \frac{1}{\sqrt{\delta}} \frac{\partial}{\partial x^i} (\sqrt{\delta} a^i)} \quad (3.3)$$

where $\delta = \det(\delta_{ij})$.

• Example (Cartesian coordinates)

$$\delta_{ij} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix} \Rightarrow \delta = 1$$

$$\Rightarrow \nabla_i a^i = \frac{\partial a^i}{\partial x^i} = \frac{\partial a^1}{\partial x^1} + \frac{\partial a^2}{\partial x^2} + \frac{\partial a^3}{\partial x^3} \quad (3.4)$$



• Relativistic version

This should involve ρ, u^ν and ∇_ν , the covariant derivative of spacetime, in order to have a 4-tensor equation. The simplest one is

$$\boxed{\nabla_\nu (\rho u^\nu) = 0} \quad (3.5)$$

For the 4-divergence we have

$$\boxed{\nabla_\nu A^\nu = \frac{1}{\sqrt{-g}} \frac{\partial}{\partial x^\nu} (\sqrt{-g} A^\nu)} \quad (3.6)$$

where $g = \det(g_{\mu\nu})$.

• 3+1 splitting at inertial frames

$$\begin{aligned} \nabla_\nu A^\nu &= \frac{1}{\sqrt{g}} \frac{\partial}{\partial x^0} (\sqrt{g} A^0) = \frac{1}{\sqrt{g}} \frac{1}{c} \frac{\partial}{\partial t} (\sqrt{g} A^0) + \frac{1}{\sqrt{g}} \frac{\partial}{\partial x^i} (\sqrt{g} A^i) \\ &= \frac{1}{c} \frac{\partial}{\partial t} (A^0) + \frac{1}{\sqrt{g}} \frac{\partial}{\partial x^i} (\sqrt{g} A^i) \end{aligned}$$

Thus, (3.5) reads

$$\begin{aligned} \frac{1}{c} \frac{\partial}{\partial t} (\rho \Gamma) + \frac{1}{\sqrt{g}} \frac{\partial}{\partial x^i} (\sqrt{g} \rho \Gamma v^i) &= 0 \\ \Rightarrow \boxed{\frac{\partial}{\partial t} (\rho \Gamma) + \frac{1}{\sqrt{g}} \frac{\partial}{\partial x^i} (\sqrt{g} \rho \Gamma v^i) = 0} & \quad (3.7) \end{aligned}$$

This can also be written in vector form as

$$\frac{\partial}{\partial t} (\rho \Gamma) + \nabla \cdot (\rho \Gamma \underline{v}) = 0.$$

3.2 Energy-momentum equation

$$\boxed{\nabla_\nu T^{\nu\mu} = 0 \quad \text{or} \quad \nabla_\nu T^\nu{}_\mu = 0} \quad (3.8)$$

where $T^{\nu\mu} = T^{\mu\nu}$ is the stress-energy-momentum tensor

• 3+1 splitting of inertial frames:

T^{00} - energy density

cT^{0i} - 3-vector of energy flux

T^{0i}/c - 3-vector of momentum density

T^{ij} - 3-stress tensor (flux of momentum)

3.3 Stress-energy-momentum of perfect gas

In the fluid frame

$$T^{00} = \varepsilon, \quad T^{0i} = T^{i0} = 0$$

$$T^{ij} = P\delta^{ij} \quad \text{- pure pressure.}$$

Thus,

$$T^{\nu\mu} = \begin{pmatrix} \varepsilon & 0 & 0 & 0 \\ 0 & & & \\ 0 & & & \\ 0 & & & \end{pmatrix} \begin{matrix} \\ (P\delta^{ij}) \\ \\ \end{matrix} \quad (3.9)$$

In general $T^{\nu\mu}$ is given by the tensor equation

$$\boxed{T^{\nu\mu} = \frac{\omega}{c^2} u^\nu u^\mu + P g^{\nu\mu}} \quad (3.10)$$

In order to verify (3.10) it is just sufficient to show

that (3.10) reduces to (3.9) in the fluid frame, where

$$u^\alpha = (c, \underline{0}) \quad g_{\alpha\beta} = \begin{pmatrix} -1 & 0 & 0 & 0 \\ 0 & & & \\ 0 & & & \\ 0 & & & \delta^{ij} \end{pmatrix}$$

Indeed

$$T^{00} = \frac{\omega}{c^2} u^0 u^0 + P g^{00} = \frac{\epsilon + P}{c^2} c^2 - P = \epsilon$$

$$T^{0i} = \frac{\omega}{c^2} u^0 u^i + P g^{0i} = 0 + 0 = 0$$

$$T^{ij} = \frac{\omega}{c^2} u^i u^j + P g^{ij} = 0 + P \delta^{ij} = P \delta^{ij}$$

3.4 3+1 splitting of inertial frames

For the divergence of symmetric tensor one has

$$\nabla_\nu T^\nu_\mu = \frac{1}{\sqrt{g}} \frac{\partial}{\partial x^\nu} (\sqrt{g} T^\nu_\mu) - \frac{1}{2} \frac{\partial g_{\alpha\beta}}{\partial x^\mu} T^{\alpha\beta} \quad (3.10)$$

For inertial frames of special relativity

$$\nabla_\nu T^\nu_0 = \frac{1}{\sqrt{g}} \frac{\partial}{\partial x^\nu} (\sqrt{g} T^\nu_0) - \frac{1}{2c} \frac{\partial g_{\alpha\beta}}{\partial t} T^{\alpha\beta}$$

Since $\partial g_{\alpha\beta} / \partial t = 0$ (static grid)

$$\begin{aligned} \nabla_\nu T^\nu_0 &= \frac{1}{\sqrt{g}} \frac{\partial}{\partial x^\nu} (\sqrt{g} T^\nu_0) = \frac{1}{\sqrt{g}} \frac{\partial}{\partial x^0} (\sqrt{g} T^0_0) + \frac{1}{\sqrt{g}} \frac{\partial}{\partial x^i} (\sqrt{g} T^i_0) \\ &= \frac{1}{c} \frac{\partial}{\partial t} (T^0_0) + \frac{1}{\sqrt{g}} \frac{\partial}{\partial x^i} (\sqrt{g} T^i_0). \end{aligned}$$

Thus, the energy equation is

$$\frac{\partial}{\partial t} (T^0_0) + \frac{1}{\sqrt{g}} \frac{\partial}{\partial x^i} (\sqrt{g} T^i_0) = 0 \quad (3.11)$$

Similarly we have the momentum equation:

$$\boxed{\frac{\partial}{\partial t}(T^0_j/c) + \frac{1}{\sqrt{g}} \frac{\partial}{\partial x^i} (\sqrt{g} T^i_j) - \frac{1}{2} \frac{\partial g_{ab}}{\partial x^j} T^{ab} = 0} \quad (3.12)$$

• If x^i are Cartesian coordinates then

$$\delta = 1, \quad \frac{\partial g_{ab}}{\partial x^j} = 0$$

$$T^0_j = T^{0j}, \quad T^i_j = T^{ij}, \quad T^0_0 = -T^{00}, \quad T^i_0 = -T^{i0}$$

and (3.11), (3.12) reduce to

$$\frac{\partial T^{00}}{\partial t} + \frac{\partial}{\partial x^i} (c T^{i0}) = 0 \quad (3.13)$$

$$\frac{\partial}{\partial t} (T^{0j}/c) + \frac{\partial}{\partial x^i} (T^{ij}) = 0 \quad (3.14)$$

Substituting T^{ab} from (3.10) one obtains

$$\boxed{\frac{\partial}{\partial t} (\omega \Gamma^2 - p) + \frac{\partial}{\partial x^i} (\omega \Gamma^2 v^i) = 0} \quad (3.15)$$

$$\boxed{\frac{\partial}{\partial t} \left(\frac{\omega \Gamma^2}{c} v^i \right) + \frac{\partial}{\partial x^i} \left(\frac{\omega \Gamma^2}{c} v^i v^j + p \delta^{ij} \right) = 0} \quad (3.16)$$

where δ^{ij} is the Kronecker's delta.

§4 Electromagnetism

4.1 Levi-Civita tensor of space

Levi-Civita symbol:

$$\epsilon^{ijk} \equiv \epsilon_{ijk} = \begin{cases} 1 & \text{if } i,j,k \text{ is even permutation of } 1,2,3 \\ -1 & \text{if } i,j,k \text{ is odd permutation of } 1,2,3 \\ 0 & \text{otherwise} \end{cases} \quad (4.1)$$

Levi-Civita tensor of space

$$\epsilon_{ijk} = \begin{cases} \sqrt{\delta} \epsilon_{ijk} & \text{for right-handed coordinates} \\ -\sqrt{\delta} \epsilon_{ijk} & \text{for left-handed coordinates} \end{cases} \quad (4.2)$$

$$\epsilon^{ijk} = \begin{cases} \frac{1}{\sqrt{\delta}} \epsilon^{ijk} & \text{for right-handed} \\ -\frac{1}{\sqrt{\delta}} \epsilon^{ijk} & \text{for left-handed} \end{cases} \quad (4.3)$$

Usage

$$1) \quad \underline{E} = \underline{B} \times \underline{V} \iff \begin{cases} E_i = \epsilon_{ijk} B^j V^k \\ E_i = \epsilon^{ijk} B_j V_k \end{cases} \quad (4.4)$$

$$2) \quad \underline{B} = \underline{\nabla} \times \underline{A} \iff \begin{cases} B^i = \epsilon^{ijk} \nabla_j A_k \\ B_i = \epsilon_{ijn} \nabla^j A^n \end{cases} \quad (4.5)$$

3) Dual vectors and anti-symmetric tensors:

$$\boxed{\begin{aligned} \text{If } F^{ij} &= \epsilon^{ijk} B_k \\ \text{then } B_k &= \frac{1}{2} \epsilon_{kij} F^{ij} \end{aligned}} \quad (4.6)$$

Notice that

$$F^{ij} = -F^{ji}$$

and has only 3 independent components, just as B^k . Hence both contain the same amount of information.

4.2 The Maxwell tensor

B and E are 3-vectors. Hence 6 components.

For a covariant formulation of electrodynamics we need a spacetime tensor.

- (i) 1-rank tensor (4-vector) \rightarrow 4 components \rightarrow too few
- (ii) 2-rank tensor \rightarrow 16 components \rightarrow too many
- (iii) 2-rank symmetric tensor \rightarrow 10 components \rightarrow still too many
- (iv) 2-rank anti-symmetric tensor \rightarrow 6 components \rightarrow exact match!

Hence introduce an anti-symmetric tensor

$$F^{\nu\mu} = -F^{\mu\nu}, \quad (4.7)$$

the Maxwell tensor, to describe the electromagnetic field.

The Lorentz 4-force acting on a charge q with 4-velocity u^ν has to be

$$\phi^\nu = \alpha q F^{\nu\mu} u_\mu \quad (4.8)$$

where α is a space-time scalar (constant)

~~4.3~~ 3+1 splitting of inertial frames

We have to split $F^{\mu\nu}$ into \underline{E} and \underline{B} at inertial frames in such a way that (4.8) give the Lorentz 3-force

$$\underline{F} = q(\underline{E} + \frac{1}{c} \underline{v} \times \underline{B}) \quad (4.9)$$

If we put

$$\boxed{E_i = F_{i0} \quad (E^i = -F^{i0})} \quad (4.10)$$

and

$$\boxed{B_i = \frac{1}{2} \epsilon_{ijk} F^{jk} \quad F^{jk} = \epsilon^{jks} B_s} \quad (4.11)$$

then we recover (4.9) from (4.8) if $\alpha = \frac{1}{c}$.

Thus.

$$\boxed{\phi^\nu = \frac{q}{c} F^{\nu\mu} u_\mu} \quad (4.12)$$

Exercise Derive (4.9) from (4.11) (4.12).

$$\phi_0 = \frac{dP_0}{d\tau} \Rightarrow \phi_i = \frac{dP_i}{d\tau} = \Gamma \frac{dP_i}{dt} = \Gamma f_i \quad (*)$$

$$(4.12) \Rightarrow \phi_i = \frac{q}{c} (F_{i0} u^0 + F_{ij} u^j) =$$

$$= \frac{q u^0}{c} (F_{i0} + F_{ij} \frac{u^j}{u^0}) = \frac{q u^0}{c} (E_i + \frac{1}{c} \epsilon_{ijk} B^k v^j) =$$

$$= q \Gamma (E_i + \frac{1}{c} \epsilon_{ijk} v^j B^k)$$

$$\Rightarrow f_i = q (E_i + \frac{1}{c} \epsilon_{ijk} v^j B^k)$$

$$\text{or } \underline{f} = q (\underline{E} + \frac{1}{c} \underline{v} \times \underline{B}) \quad \square$$

According to (4.10) and (4.11) the 3+1 splitting is

$$F^{\mu\nu} = \begin{pmatrix} 0 & E_1 & E_2 & E_3 \\ -E_1 & 0 & B_3 & -B_2 \\ -E_2 & -B_3 & 0 & B_1 \\ -E_3 & -B_1 & B_2 & 0 \end{pmatrix} \quad (4.13)$$

4.3 The Faraday tensor

4-dimensional Levi-Civita symbol:

$$e^{\alpha\beta\gamma\delta} = \epsilon_{\alpha\beta\gamma\delta} = \begin{cases} 1 & \text{if } \alpha, \beta, \gamma, \delta \text{ is an even permutation of } 0, 1, 2, 3 \\ -1 & \text{if } \alpha, \beta, \gamma, \delta \text{ is an odd permutation of } 0, 1, 2, 3 \\ 0 & \text{otherwise} \end{cases} \quad (4.14)$$

Levi-Civita tensor of space-time

$$e_{\alpha\beta\gamma\delta} = \begin{cases} \sqrt{-g} \epsilon_{\alpha\beta\gamma\delta} & \text{in right-handed coordinates} \\ -\sqrt{-g} \epsilon_{\alpha\beta\gamma\delta} & \text{in left-handed} \end{cases} \quad (4.15)$$

From this it follows that

$$e^{\alpha\beta\gamma\delta} = \begin{cases} -\frac{1}{\sqrt{-g}} \epsilon^{\alpha\beta\gamma\delta} & \text{in right-handed systems} \\ \frac{1}{\sqrt{-g}} \epsilon^{\alpha\beta\gamma\delta} & \text{in left-handed} \end{cases} \quad (4.16)$$

The Faraday tensor

$$\boxed{*F^{\mu\nu} = \frac{1}{2} e^{\mu\nu\alpha\beta} F_{\alpha\beta}} \quad (4.17)$$

From this we have

$$\boxed{*F^{\mu\nu} = -*F^{\nu\mu}} \quad (4.18)$$

$*F^{\mu\nu}$ is also anti-symmetric and hence contains the same amount of information as $F^{\mu\nu}$. Hence, it can also describe the electromagnetic field.

From (4.17) it follows that

(4.6)

$$\boxed{F^{\mu\nu} = -\frac{1}{2} \epsilon^{\mu\nu\alpha\beta} *F_{\alpha\beta}} \quad (4.19)$$

• 3+1 splitting in inertial frames

$$\boxed{B_i = *F_{0i}} \quad (4.20)$$

$$\boxed{E^i = \frac{1}{2} \epsilon^{ijk} *F_{jk}} \quad (4.21)$$

4.4 Maxwell equations

$$\boxed{\nabla_\beta *F^{\alpha\beta} = 0} \quad \text{- Faraday group} \quad (4.22)$$

$$\boxed{\nabla_\beta F^{\alpha\beta} = \frac{4\pi}{c} I^\alpha} \quad \text{- Maxwell group} \quad (4.23)$$

where I^α is the electric current 4-vector

• 3+1 splitting of inertial frames

$$I^\alpha = (c\rho, \underline{j}) \quad (4.24)$$

ρ - electric charge density

\underline{j} - electric current density

$$(4.22) \Rightarrow \int \underline{\nabla} \cdot \underline{B} = 0 \quad (4.25)$$

$$\int \frac{1}{c} \frac{\partial B}{\partial t} + \underline{\nabla} \times \underline{E} = 0 \quad (4.26)$$

$$(4.23) \Rightarrow \int \underline{\nabla} \cdot \underline{E} = 4\pi \rho \quad (4.27)$$

$$\int -\frac{1}{c} \frac{\partial E}{\partial t} + \underline{\nabla} \times \underline{B} = \frac{4\pi}{c} \underline{j} \quad (4.28)$$

These are obtained using the result

$$\nabla_{\beta} F^{\alpha\beta} = \frac{1}{\sqrt{-g}} \frac{\partial}{\partial x^{\beta}} (\sqrt{-g} F^{\alpha\beta}) \quad (4.29)$$

for any anti-symmetric tensor.

• Exercise (Derive (4.25) and (4.26))

(4.24) reads

$$\frac{1}{\sqrt{-g}} \frac{\partial}{\partial x^{\beta}} (\sqrt{-g} \star F^{\alpha\beta}) = 0$$

$$\Rightarrow \frac{1}{\sqrt{g}} \frac{\partial}{\partial t} (\sqrt{g} \star F^{\alpha 0}) = 0 \Rightarrow \frac{\partial}{\partial x^{\beta}} (\sqrt{g} \star F^{\alpha\beta}) = 0$$

$$\Rightarrow \frac{1}{c} \frac{\partial}{\partial t} (\sqrt{g} \star F^{\alpha 0}) + \frac{\partial}{\partial x^i} (\sqrt{g} \star F^{\alpha i}) = 0$$

$$\Rightarrow \frac{1}{c} \frac{\partial}{\partial t} (\star F^{\alpha 0}) + \frac{1}{\sqrt{g}} \frac{\partial}{\partial x^i} (\sqrt{g} \star F^{\alpha i}) = 0$$

$\alpha=0$

$$\frac{1}{\sqrt{g}} \frac{\partial}{\partial x^i} (\sqrt{g} \star F^{0i}) = 0$$

$$\Rightarrow \frac{1}{\sqrt{g}} \frac{\partial}{\partial x^i} (\sqrt{g} B^i) = 0 \Rightarrow \underline{\nabla} \cdot \underline{B} = 0 \quad //$$

$$\underline{\alpha = j}$$

$$\frac{1}{c} \frac{\partial}{\partial t} {}^*F^{j0} + \frac{1}{\sqrt{\gamma}} \frac{\partial}{\partial x^i} (\sqrt{\gamma} {}^*F^{ji}) = 0 \quad (*)$$

$${}^*F^{j0} = -{}^*F^{0j} = B^j$$

$${}^*F^{ji} = \epsilon^{jik} E_k = \frac{1}{\sqrt{\gamma}} \epsilon^{jik} E_k$$

Thus (*) reads

$$\frac{1}{c} \frac{\partial}{\partial t} (B^j) + \frac{1}{\sqrt{\gamma}} \frac{\partial}{\partial x^i} (\epsilon^{jik} E_k) = 0$$

$$\Rightarrow \frac{1}{c} \frac{\partial}{\partial t} (B^j) + \frac{1}{\sqrt{\gamma}} \epsilon^{jik} \frac{\partial}{\partial x^i} (E_k) = 0$$

$$\Rightarrow \frac{1}{c} \frac{\partial}{\partial t} (B^j) + \epsilon^{jik} \frac{\partial}{\partial x^i} E_k = 0$$

$$\Rightarrow \frac{1}{c} \frac{\partial \underline{B}}{\partial t} + \underline{\nabla} \times \underline{E} = 0 \quad \ll$$

□

4.5 Stress-energy-momentum tensor

$$\boxed{T^{\mu\nu} = \frac{1}{4\pi} (F^{\mu\alpha} F^{\nu}_{\alpha} - \frac{1}{4} (F_{\alpha\beta} F^{\alpha\beta}) g^{\mu\nu})} \quad (4.30)$$

• 3+1 splitting of inertial frames

$$T^{00} = \frac{1}{8\pi} (E^2 + B^2), \quad T^{0i} = \frac{1}{4\pi} \epsilon^{ijk} E_j B_k$$

$$T^{ij} = \frac{1}{4\pi} \left[-(E^i E^j + B^i B^j) + \frac{1}{2} \gamma^{ij} (E^2 + B^2) \right] \quad (4.31)$$

4.6 Ohm's Law

The simplest 3+1 form in the fluid frame is

$$\boxed{\underline{j} = \sigma \underline{E}} \quad (4.32)$$

where σ is the conductivity.

The corresponding covariant form is

$$\boxed{I_\nu = \frac{\sigma}{c} F_{\nu\mu} u^\mu + q_0 u_\nu} \quad (4.33)$$

where

$$q_0 = -\frac{I_0 u^0}{c^2} \quad (4.34)$$

is the electric charge density in the fluid frame.

• Exercise (Derive (4.32) from (4.33))

In the fluid frame $u^0 = (c, \underline{0})$, $u_\nu = (-c, \underline{0})$,

$I_0 = (-c q_0, \underline{j})$, $F_{\nu\mu} u^\mu = F_{\nu 0} c = (0, \underline{E} c)$

Thus, (4.33) reads

$$(-c q_0, \underline{j}) = \frac{\sigma}{c} (0, \underline{E} c) + q_0 (-c, \underline{0})$$

$$\Rightarrow (-c q_0, \underline{j}) = (-c q_0, \sigma \underline{E})$$

$$\Rightarrow \left. \begin{array}{l} q = q_0 \\ \underline{j} = \sigma \underline{E} \end{array} \right\}$$

$$\square$$

• The perfect conductivity limit ($\sigma \rightarrow \infty$)

It remains finite only if $F_{\nu\mu} u^\mu \rightarrow 0$
Thus, in this limit

$$\boxed{F_{\nu\mu} u^\mu = 0} \quad (4.35)$$

In the fluid frame (4.35) reads

$$F_{\nu 0} = 0 \Rightarrow F_{i0} = 0 \Rightarrow \boxed{\underline{E} = 0} \quad (4.36)$$

General 3+1 splitting of inertial frames

$$F_{\nu\mu} u^\mu = 0 \begin{cases} \rightarrow F_{0\mu} u^\mu = 0 \\ \rightarrow F_{i\mu} u^\mu = 0 \end{cases}$$

$$F_{0\mu} u^\mu = 0 \Rightarrow F_{0i} u^i = 0 \Rightarrow E_i v^i = 0 \Rightarrow \boxed{(\underline{E} \cdot \underline{v}) = 0} \quad (4.37)$$

$$\begin{aligned} F_{i\mu} u^\mu &= F_{i0} u^0 + F_{ij} u^j = 0 \\ \Rightarrow E_i \Gamma + \epsilon_{ijk} B^k \Gamma v^j &= 0 \\ \Rightarrow E_i + \frac{1}{c} \epsilon_{ijk} v^j B^k &= 0 \\ \Rightarrow \boxed{\underline{E} + \frac{1}{c} \underline{v} \times \underline{B} = 0} &\quad (4.38) \end{aligned}$$

- the same result as in non-relativistic MHD

5.5 Relativistic MHD

5.1 Resistive RMHD

- The energy-momentum equation

$$\boxed{\nabla_\mu (T_{(m)}^{\nu\mu} + T_{(e)}^{\nu\mu}) = 0} \quad (5.1)$$

$T_{(m)}^{\nu\mu}$ - tensor of matter (eq. 3.10)

$T_{(e)}^{\nu\mu}$ - tensor of electromagnetic field (eq 4.30)

Using the Maxwell equations

$$\nabla_\mu (T_{(m)}^{\nu\mu}) = -\nabla_\mu T_{(e)}^{\nu\mu} = \frac{1}{c} F^{\mu\sigma} I_\sigma \quad (5.2)$$

Here
$$\phi^\mu = \frac{1}{c} F^{\mu\sigma} I_\sigma \quad (5.3)$$

is the MHD version of the Lorentz 4-force.
(Compare with eq. 4.12)

- The continuity equation

$$\boxed{\nabla_\sigma \rho u^\sigma = 0} \quad (5.4)$$

- Maxwell's equations

$$\boxed{\nabla_\beta \star F^{\alpha\beta} = 0} \quad (5.5)$$

$$\boxed{\nabla_\beta F^{\alpha\beta} = I^\alpha} \quad (5.6)$$

• Ohm's Law

$$\boxed{I_0 = \frac{6}{c} F_{\nu\mu} u^\mu + g_0 u_\nu} \quad (5.7)$$

$$g_0 = -\frac{1}{c^2} (I_0 u^\nu)$$

5.2 Ideal RMHD

$$6 = \infty \Rightarrow \boxed{F_{\nu\mu} u^\mu = 0} \quad (5.8)$$

This equation replaces the Maxwell equation (5.6)

5.3 Ideal RMHD in the Lichnerowicz Formulation

Introduce 4-vectors of electric

$$\boxed{e_\nu = \frac{1}{c} F_{\nu\mu} u^\mu} \quad (5.9)$$

and magnetic fields

$$\boxed{b_\nu = -\frac{1}{c} *F_{\nu\mu} u^\mu} \quad (5.10)$$

From this we have

$$e_\nu u^\nu = 0 \quad (5.11)$$

$$b_\nu u^\nu = 0 \quad (5.12)$$

- 3+1 splitting of e_ν and b_ν in the fluid frame

$$u^\mu = (c, \underline{0}) \Rightarrow e_\nu = F_{\nu 0} = (0, \underline{E}) \quad (5.13)$$

$$b_\nu = -\frac{1}{c} {}^*F_{\nu 0} c = (0, \underline{B}) \quad (5.14)$$

- $F_{\nu\mu}$ and ${}^*F_{\nu\mu}$ in terms of b_ν, e_ν and u_ν

$$F_{\nu\mu} = \frac{1}{c} [u_\nu e_\mu - e_\nu u_\mu + e_{\nu\mu\alpha\beta} u^\alpha b^\beta] \quad (5.15)$$

$${}^*F_{\nu\mu} = \frac{1}{c} [b_\nu u_\mu - u_\nu b_\mu + e_{\nu\mu\alpha\beta} u^\alpha e^\beta] \quad (5.16)$$

- ~~RMHD equations~~
Ideal RMHD equations

Since in the fluid frame $\underline{E} = 0$ we have $e_\nu = 0$, see also (5.9) and (5.8).

Then

$$F_{\nu\mu} = \frac{1}{c} e_{\nu\mu\alpha\beta} u^\alpha b^\beta \quad (5.17)$$

$${}^*F_{\nu\mu} = \frac{1}{c} [b_\nu u_\mu - u_\nu b_\mu] \quad (5.18)$$

Combining (5.17) with (4.30) one obtains

$$T^{\nu\mu}(\mathbf{e}) = \frac{1}{4\pi} \left[\frac{b^2}{c^2} u^\nu u^\mu - b^\nu b^\mu + \frac{b^2}{2} g^{\nu\mu} \right] \quad (5.19)$$

(5.4)

Given (5.18) the Faraday equation becomes

$$\nabla_{\mu} (b^{\nu} u^{\mu} - b^{\mu} u^{\nu}) = 0 \quad (5.20)$$

Given (5.19) and (5.10) the energy-momentum equation becomes

$$\nabla_{\mu} \left[\left(\omega + \frac{b^2}{4\pi} \right) \frac{u^{\nu} u^{\mu}}{c^2} - \frac{b^{\nu} b^{\mu}}{4\pi} + \left(p + \frac{b^2}{4\pi} \right) g^{\nu\mu} \right] = 0 \quad (5.21)$$

The continuity equation is as before

$$\nabla_{\alpha} (\rho u^{\alpha}) = 0 \quad (5.23)$$

These PDEs are complemented with

$$\omega = \omega(p, \rho) \quad - \text{EOS}$$

$$b_{\nu} u^{\nu} = 0$$

$$u_{\nu} u^{\nu} = -c^2$$

Notice that the magnetic field contributes not only to stress-tensor but also to fluid inertia via the term $b^2/4\pi$ in

$$\left(\omega + \frac{b^2}{4\pi} \right) \frac{u^{\nu} u^{\mu}}{c^2} \quad \text{in (5.21)}$$