Microwave Kinetic Inductance Detectors for visible to near infrared astronomy



All of the times

mazinlab.org

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KIDs: The Next Generation - workshop DIAS Sept. 2017



Outline

- MKIDs for single photon detection
- Fabrication of kilopixels MKIDs arrays
- New MKIDs development



Single photon detector



- Single photon counting with ~ 100 microsec timing
- Energy resolving R~10



Microwave Kinetic Inductance Detectors



We use a microlens array to improve effective fill factor to ~92%



Multiplexing



10 to 20k pixels arrays



TiN : Strongly non uniform in nitrogen content over a wafer – poor control of f_0 position \rightarrow Looking for an alternative material





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Platinum Silicide: PtSi

- Resistivity: 50 $\mu\Omega$.cm
- T_c ≈ 940 mK
- We aim for 60 nm films with ~10 pH/sq inductance

Deposition of Pt and Si on sapphire substrate + anneal in-situ



Which material for the MKIDs?

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 \rightarrow Qi of 1 millions!! (on single layer test device)





TiN vs. PtSi





1) PtSi Resonator Outline
PtSi deposition and annealing + W done in-situ using AJA ATC sputter system.
ICP etching of the W (SF6) + PtSi (Cl2 + CF4 + Ar)



W

PtSi

1) PtSi Resonator Outline

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2) Nb Feedline and Ground Plane (lift-off)







3) SiO2 insulating pads

- Insulating pads over the feedline used to connect ground plane segments
- Reactive sputtering of SiO2 (alternative recently tried : aSi:H – similar performances but much easiest to deposit! 3 min vs. 3 hours!)







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4) Nb Crossovers and Coupling Bars

 Crossovers to connect ground plane segments and connections of the feedline to couplers







5) Gold Bond Pads

 Gold bond pads used for gold wire bonding (reduce heat excess)







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6) PtSi Resonator Etch

- W and PtSi Etched with ICP etcher
- Protect layer removed in heated H2O2





Dark-speckle Near-IR Energy-resolved Superconducting Spectrophotometer (DARKNESS)

Large-Format Platinum Silicide Microwave Kinetic Inductance Detectors for Optical to Near-IR Astronomy *Paul Szypryt, submitted to Optic Express (last week)*





- Qi Lower than expected (~80,000), R~8 at 1 micron
- Great sensitivity to photons in instrument's wavelength band
- Low yield due to roll off of Qi toward higher frequency
- On par or better than the best 10,000 pixel TiN arrays



DARKNESS

DARKNESS is commissioned and science runs are ongoing... (Palomar)

DARKNESS: A Microwave Kinetic Inductance Detector Integral Field Spectrograph for high-contrast Astronomy - Seth R. Meeker, submitted to PASP (last week)









MKID Exoplanet Camera (MEC)

- 20,440 (140x146) pixels split between 10 feedlines (14x146)
- 150 um pixel pitch
- 22x22 mm imaging area
- 20 ROACH2 readout boards

Improvements from DARKNESS array:

- 500 nm inductor gaps as opposed to 300 nm gaps used in DARKNESS
- 1 crossover every 10 pixels ($\sim\lambda/8$)
- Optimized Sonnet simulations for better control of designed resonator placing and 2 MHz spacing (capacitor shrinking)

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MEC results





- Our best array has Qi values of ~ 100,000
- Less variation of Qi with frequency, no significant dips
- Energy resolution of 8 at 1 micron
- Pixel yield approaching 90% (fitted resonators). *Still need more work to determine how many of those are actually photosensitive*
- Generally very good sensitivity to photons in the 700-1400 nm band



MEC results







Laminated NbTi on Kapton Microstrip Cables for Flexible Sub-Kelvin RF Electronics – A. B. Walter in preparation



- Parallel plate capacitor MKIDs
- Low Tc material



Classical LEKID design

High meandered inductance



Small interdigitated capacitance

$\rm f_0$ dominated by L

$$f_0 \propto \frac{1}{\sqrt{LC}}$$



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Parallel plate design



Major change = two large parallel plate capacitors and small square inductance

 f_0 dominated by C





- Maximize the readout power before nonlinear effects thanks to the low current density in the wide inductor (30umx30um)
- Improvement in signal-to-noise ratio due to the saturation of TLS (drive at high power)
- Inductor geometrically more uniform than a classical meandered inductor → increase energy resolution



$$S_{TLS}(\nu) = \kappa(\nu, \omega, T) \times \frac{\int_{V_{TLS}} \left| \vec{E}(\vec{r}) \right|^3 d^3 \vec{r}}{4 \left(\int_{V} \left| \mathcal{E}(\vec{r}) \vec{E}(\vec{r}) \right|^2 d^3 \vec{r} \right)^2}$$





For a parallel plate capacitor: $V_{TLS} = V$

$$S_{TLS} \propto \frac{E^3 V}{4\varepsilon^2 E^4 V^2} \propto \frac{1}{\varepsilon^2 E V}$$

Lower TLS noise by:

- Using a high ϵ material
- Maximizing the electric field in the capacitor (driving the MKID at high power)
- Making the volume of the capacitor as large as possible



Micro-fabrication process



Sputtering of Pt and Si + annealing at 300C = 55nm thick PtSi film \rightarrow Patterning of the inductor, the coupling tie and the first side of the parallel plate capacitor



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Atomic layer deposition of 10 nm of Al_2O_3 over the entire wafer (thickness uniformity of 98% over a 4inch wafer)



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Sputtering of 80 nm of Nb \rightarrow CPW feedline and second side of the capacitor



Characterization



Parallel plate MKIDs resonate

PIER 521 LGg Mag 500.0mdm/ Ref 2.000dm 1.000 51 3.7274813 GHZ 0.5185 GHZ 2.000 50.000 0.000 0.000 2.000 0.000 0.000 0.000 1.000 0.000 0.000 0.000 1.000 0.000 0.000 0.000 1.000 0.000 0.000 0.000 1.000 0.000 0.000 0.000 1.000 0.000 0.000 0.000 1.000 0.000 0.000 0.000 1.000 0.000 0.000 0.000 1.000 0.000 0.000 0.000

Test device:

- 2 feedlines (different inductor dimensions)
- 18 resonators, 3x6 centered on 4 GHz, 6 GHz and 8 GHz

- Resonances located around their design frequencies
- 13 resonances out of 18 were identified
- Few dB deep, best resonances Qi \approx 35 000 40 000



Characterization



Try next: crystalline Al2O3 annealing / ebeam deposition (pinholes?), new tri-layer insitu design Hf/HfO2/Hf



Low Tc material

Why lower
$$T_C$$
?
• Sensitivity $\propto \frac{1}{T_C^2}$
• Energy resolution $R \propto \sqrt{\frac{1}{T_C}}$



Low Tc material



Why Hafnium?

- elemental material (easy to deposit)
- T_c ~ 400mK (Δ~eV)
- High normal state resistivity

Good uniformity over a 3inch substrate Sheet resistance (ohm/sq), Hf flim / Ox Si, actual measurements in red #B161222, 1.2 min = 50 nm, 500 Watts/ 150 mm target, 2.5 mT



~ 5% variation (pretty much the same than our PtSi films)



Resonator characterization

We fabricated several Hf test devices:

- Different substrate orientation (a-place, c-place, r-plane sapphire)
- Various sputtering parameter (try to further reduce T_c)
- Annealing of the Hf in-situ

125 nm of Hf sputtered on a-plane sapphire gave good preliminary results!

- Qi up to 500 000
- R~9@808nm
- First τ_{qp} on Hf ever measured ~ 30 μ sec





Any correlation between crystal structures and performances?

\rightarrow It seems that the performances of our test device are substrate dependent



We saw :

- Nothing on C-plane
- low Qi resonators (~10k) on R-plane
- High Qi (~100k) on A-plane

Hard to find a correlation from XRD data? Need to identify theses peaks



Hafnium:

- Try to further reduce Tc with reducing the stress in the films
- Improving uniformity?
- Improve heat sinking (very low temperature measurement! Tc/8 ~ 56mK)

Other materials?



Conclusions

- We make large arrays of single photon detectors
- We achieved high performances (Qi~100 000, R~8)



DARKNESS 10 000 pixels



MEC 20 000 pixels



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 Development of parallel plates MKIDs : We demonstrate high readout power by a factor of 4 compared to classical MKIDs





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 Hafnium resonators are promising! We already have good Qi (>100 000) and energy resolution (R~9 @ 808nm) after only a few tests

Thanks!



Uniformity

- Measured sheet resistance across full 4" wafers. Early estimates show PtSi to be roughly an order of magnitude more uniform than TiN.
- For 1K T_C films, PtSi is annealed to it's thermally stable stoichiometry, whereas TiN requires precise control of the N₂ flow rate during Ti sputtering in a region where the TC is very sensitive to this Ti/N ratio.



% Variation in Sheet Resistance from Center

Szypryt et al. 2016

PtSi Pulses



- Measured quasiparticle lifetimes of 30-40 us.
- Fairly flat energy resolution of R=8 across observable wavelength band.

Szypryt et al. 2016

Quantun % Transmittance PtSi TIN % Reflectance % Absorption Wavelength (nm)

Szypryt et al. 2016

• Shaded region represents wavelength band of DARKNESS instrument.